

CMAS(s) & CMENOS?!

A road-trip to the galaxy assembly bias?

Or

A SAM perspective on massive galaxies

*Find the
Easter
Eggs (6)*



*Starring
MultiDark SAMs
Featuring
THE BOSS*

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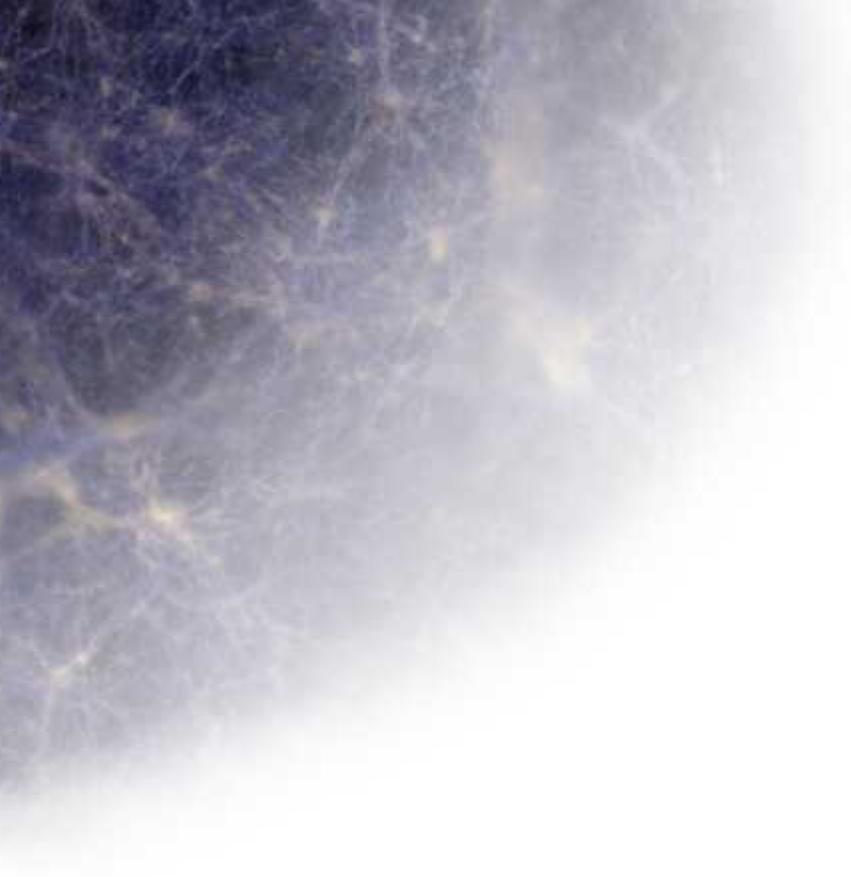


Who?!

What?!

Why?!

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SAMs

What?!

Why?!

A visualization of the cosmic web, showing a complex network of dark matter filaments and galaxy clusters in shades of blue and purple.

SAMs

LRGs/CMASS

Why?!

A visualization of the cosmic web, showing a complex network of dark matter filaments and galaxy clusters. The filaments are depicted as thin, interconnected lines, while the clusters are denser regions of matter. The overall structure is a vast, interconnected web of matter.

SAMs

LRGs/CMASS

**Galaxy Formation
Assembly bias
Cosmology**

Who?!

The MultiDark SAMs

Galacticus
(Benson+2012)

SAG
(Cora+2018)

SAGE
(Croton+2016)

$1h^{-1}$ Gpc box

Klypin+2016

Data release paper: 1710.08150
Knebe, DS, Prada et al. (2018)

<https://www.cosmosim.org/>
<http://www.skiesanduniverses.org>

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What?!

BOSS-CMASS GALAXIES

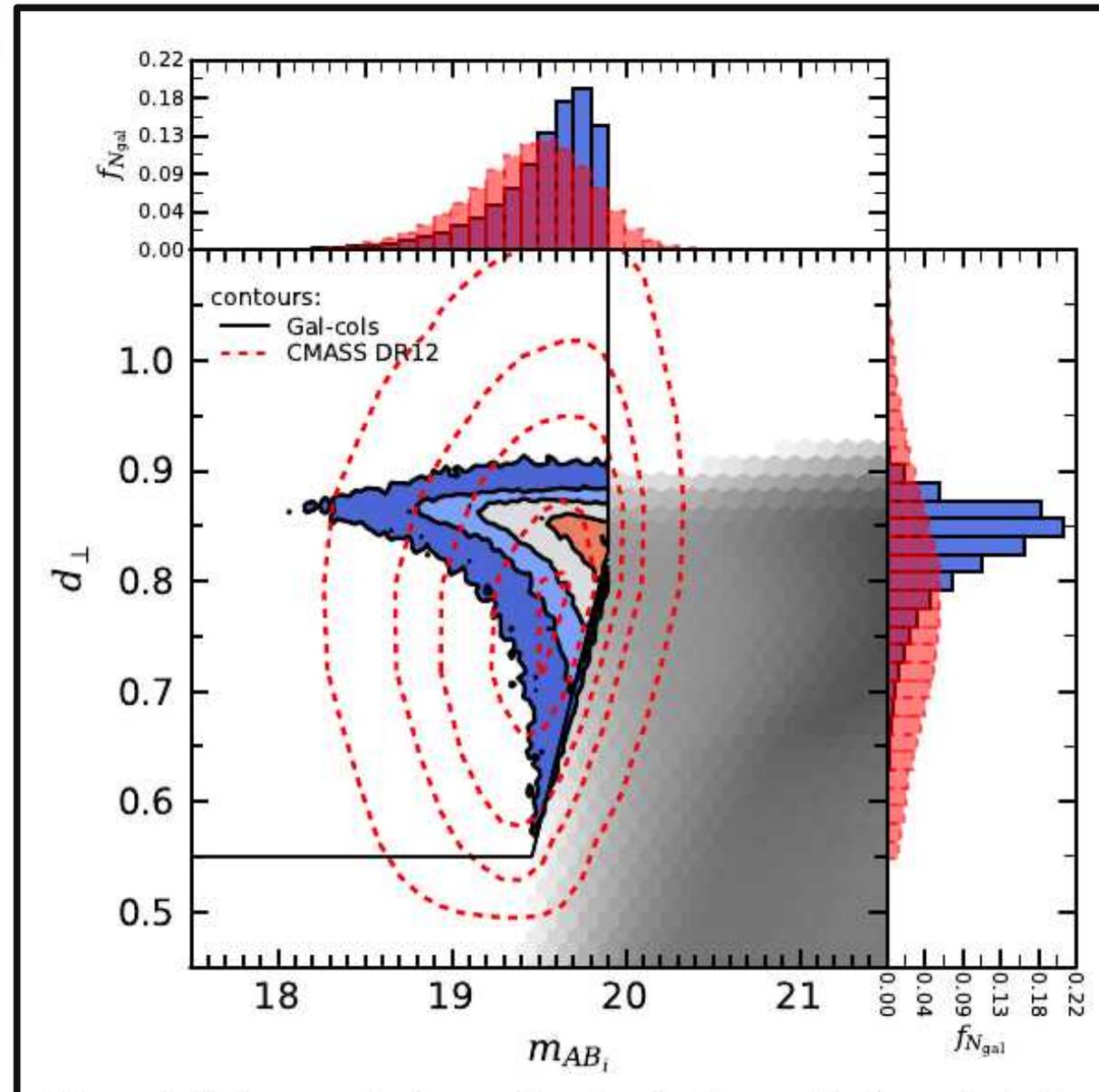
$z \sim 0.55$

“A SAM perspective
on massive
galaxies”

Stoppacher+19
arXiv: 1902.05496

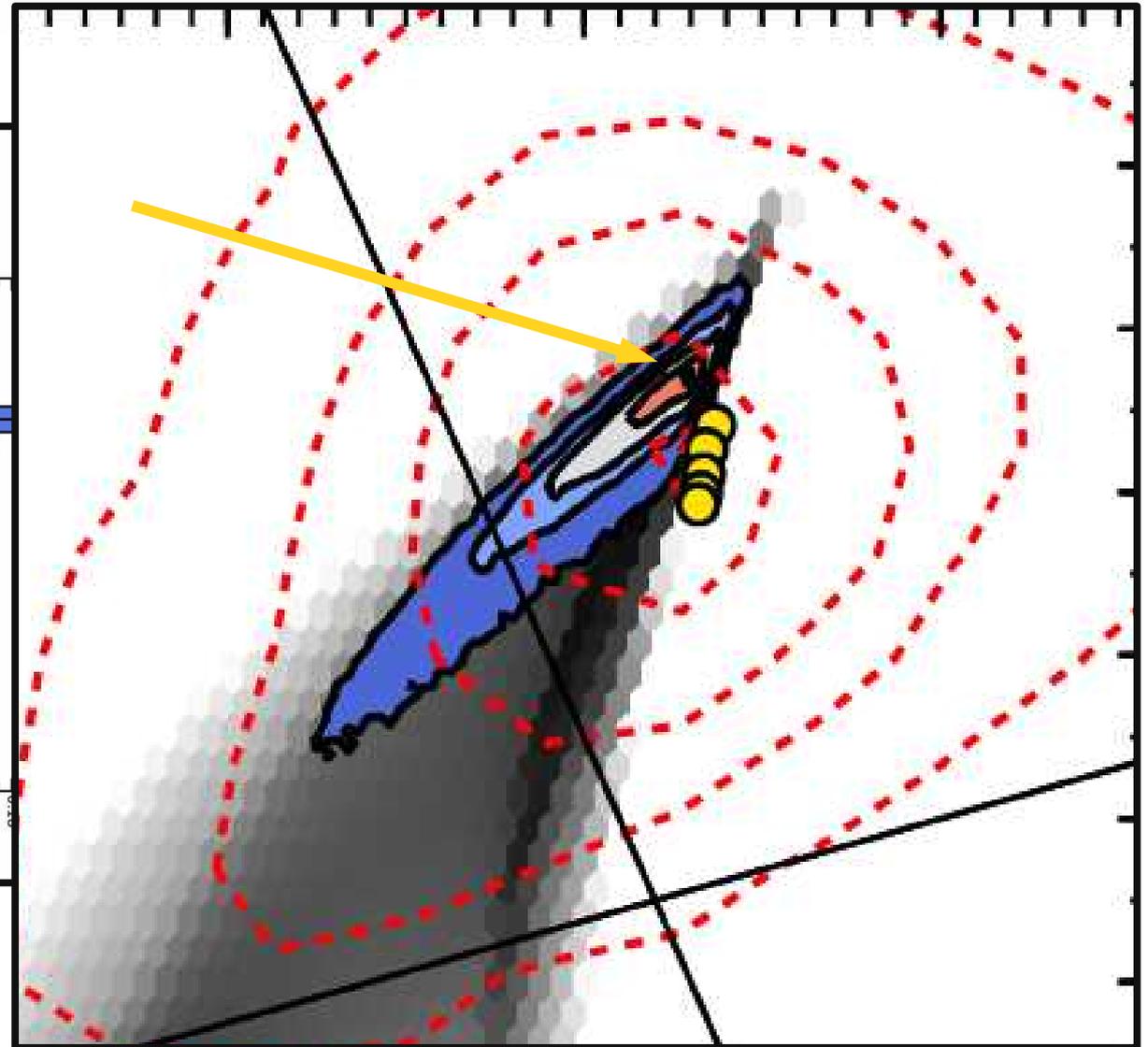
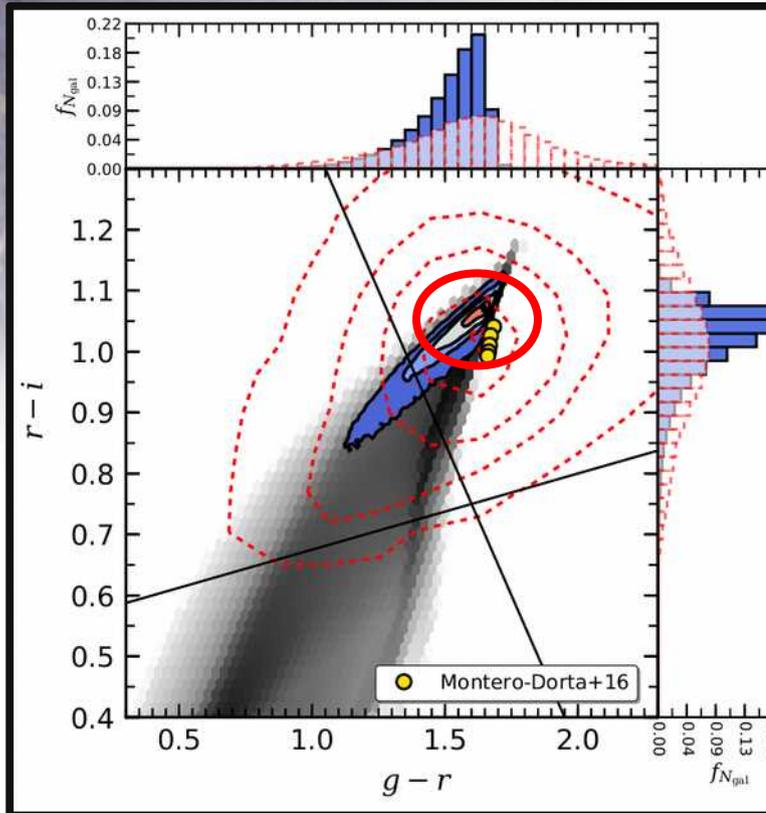
DR12 Alam+15
LSS Reid+15
Maraston+ 05 & 09

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What?!

BOSS-CMASS GALAXIES

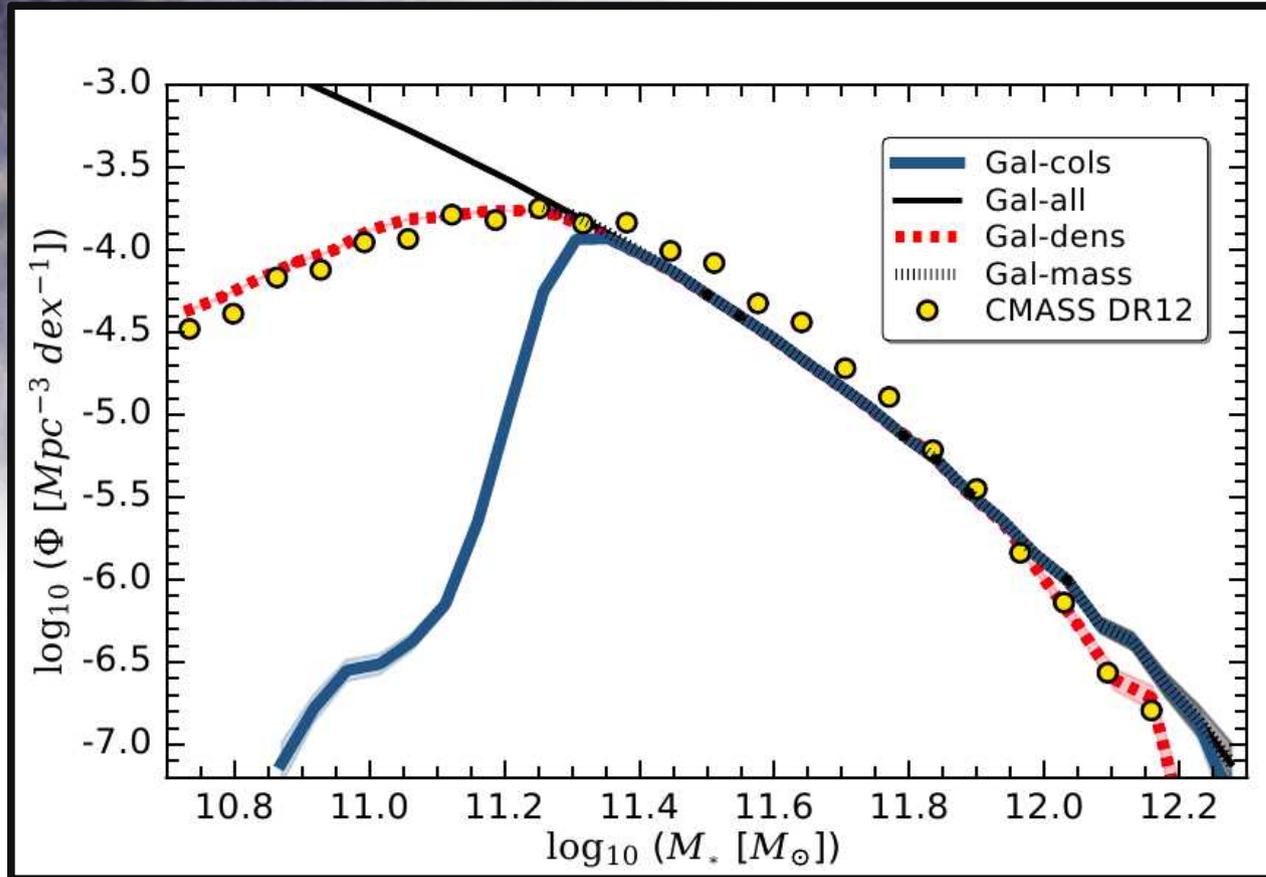


Montero-Dorta+16

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What?!

Mocking the mock



**CMASS color-cut
algorithm
“Gal-cols”**

**number density
 $3.4 \times 10^{-4} h^{-3} \text{Mpc}^{-3}$
“Gal-dens”**

**$M_{\text{star}} > 10^{11.24}$
 M_{sun}
“Gal-mass”**

Maraston+13

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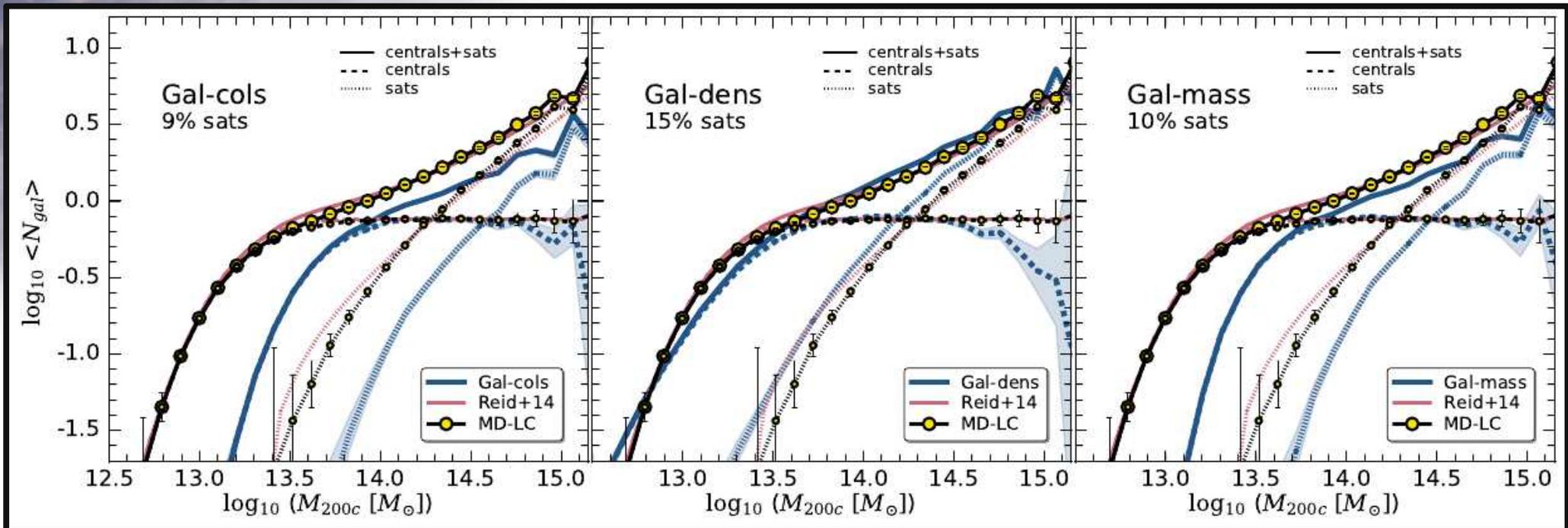
What?!

Halo Occupation Distribution

Satellites: 9%

15%

10%



CMASS color-cut
algorithm
“Gal-cols”

number density
 $3.4 \times 10^{-4} h^{-3} \text{Mpc}^{-3}$
“Gal-dens”

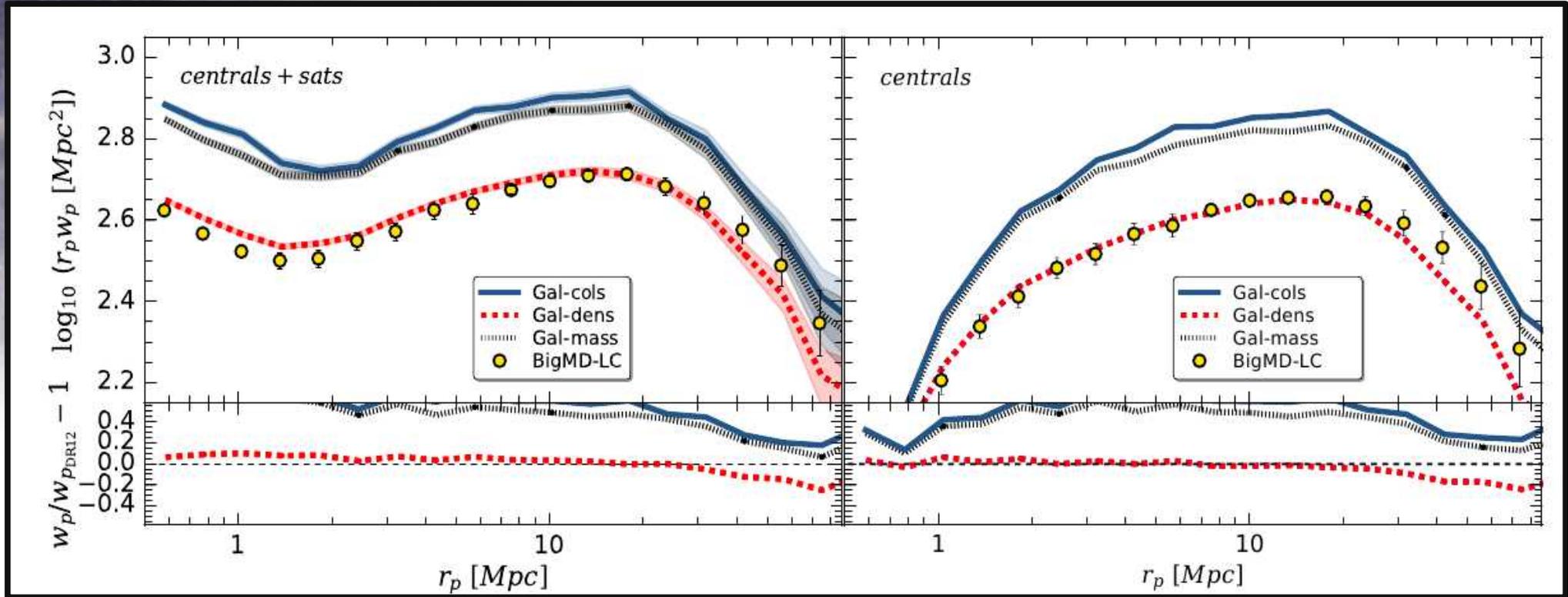
$M_{\text{star}} > 10^{11.24}$
 M_{sun}
“Gal-mass”

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Rodriguez-Torres+16

What?!

2-point Correlation Functions



CMASS color-cut
algorithm
“Gal-cols”

number density
 $3.4 \times 10^{-4} h^{-3} \text{Mpc}^{-3}$
“Gal-dens”

$M_{\text{star}} > 10^{11.24}$
Msun
“Gal-mass”

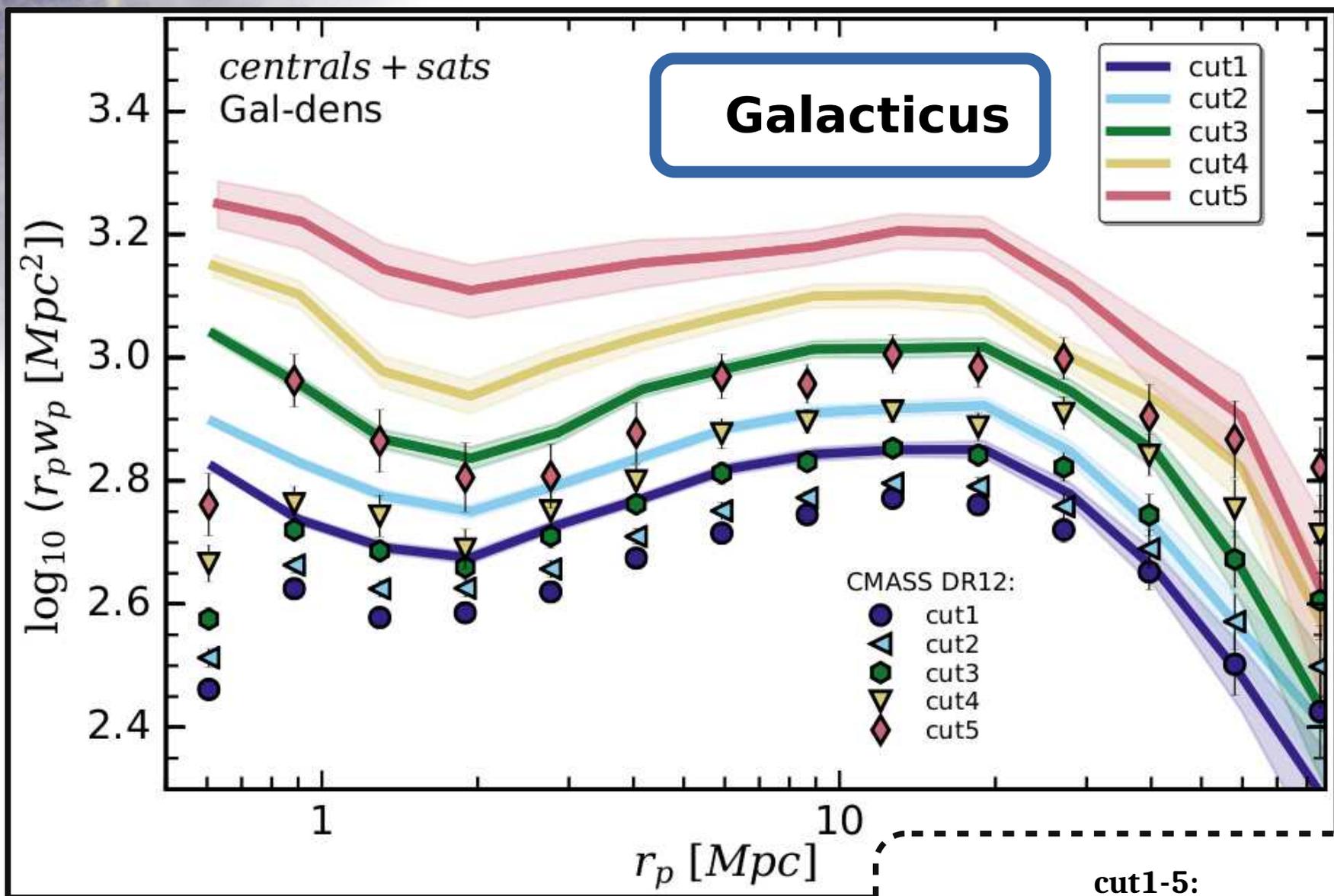
CORRFUNC
Shinha+17

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<https://pypi.python.org/pypi/Corrfunc>
<https://manodeep.github.io/Corrfunc/>

What?!

2-point CFs Stellar Mass Cuts

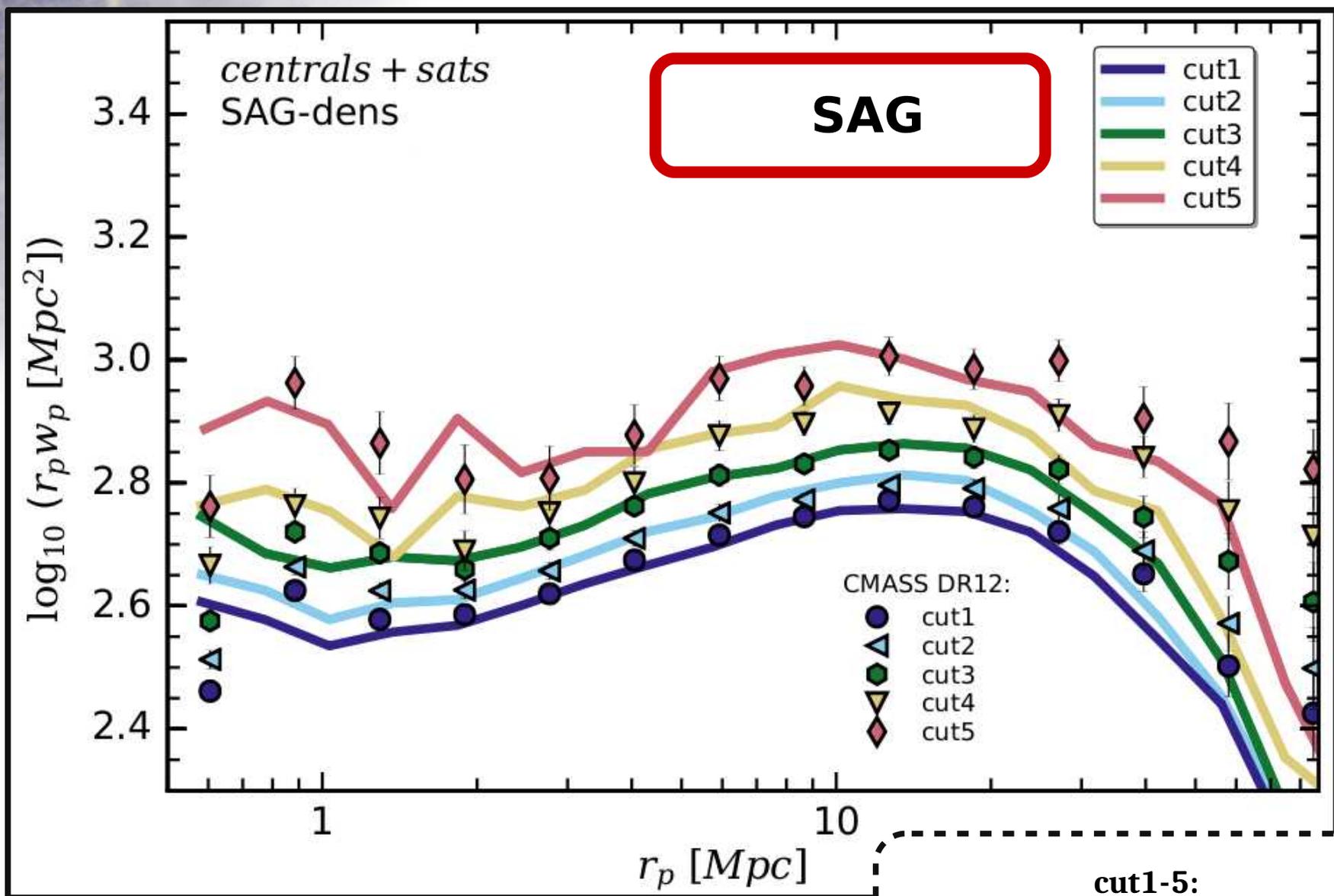


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cut1-5:
log10(Mstar [Msun]):
11.21, 11.31, 11.41, 11.51, 11.61

What?!

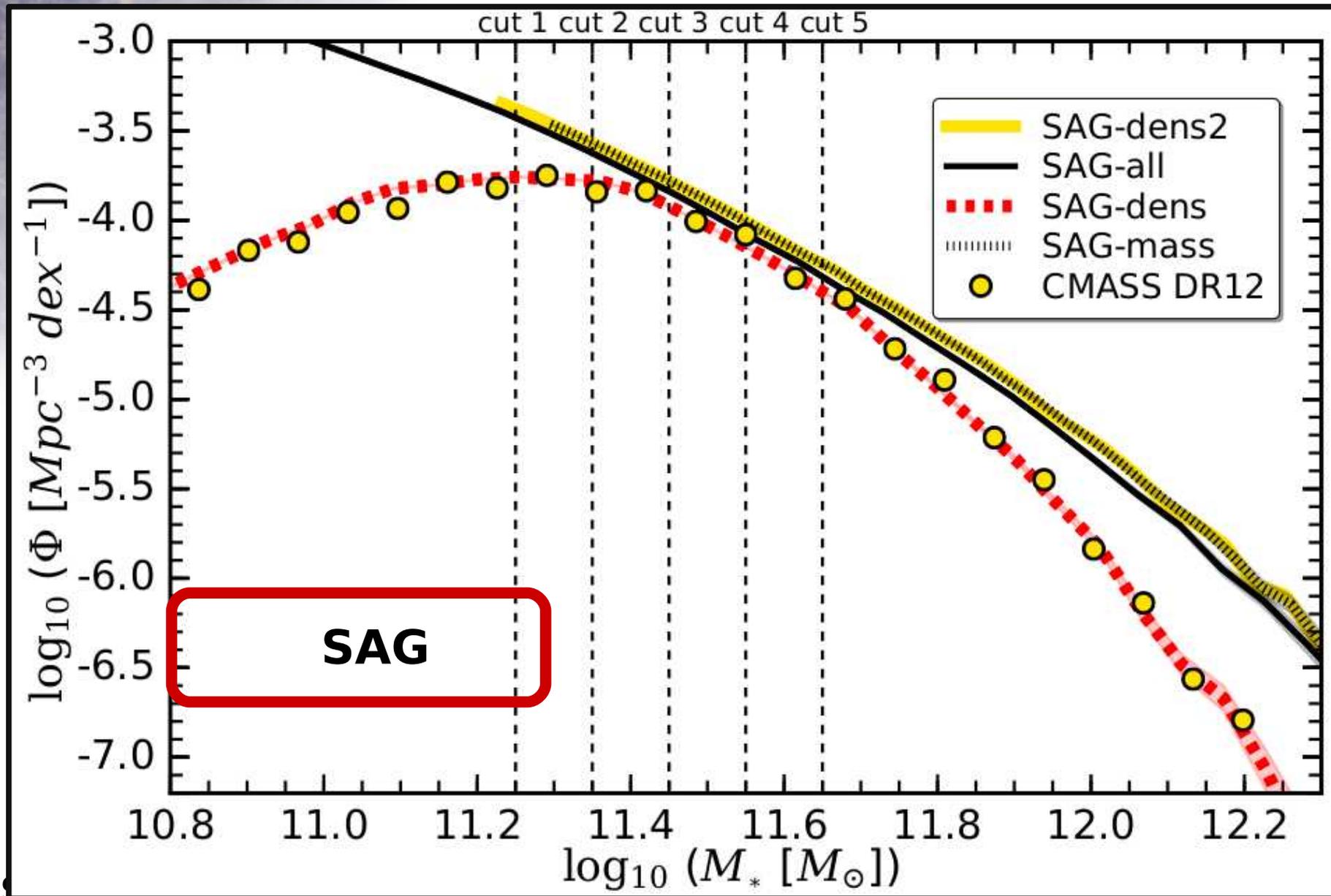
2-point CFs Stellar Mass Cuts



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What?!

Stellar Mass Function



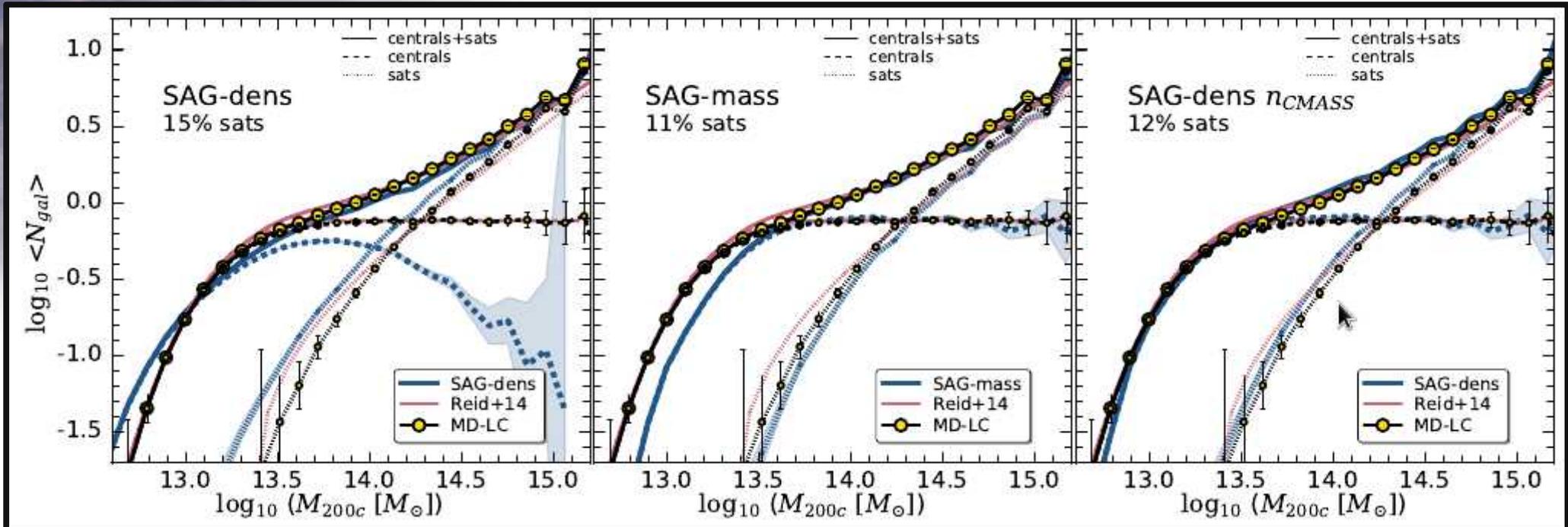
What?!

Halo Occupation Distribution

Satellites: 15%

11%

12%



Down sampled
 $3.4 \times 10^{-4} h^{-3} \text{Mpc}^{-3}$
"SAG-dens"

$M_{star} > 10^{11.24}$
Msun
"SAG-mass"

Number density
 $3.4 \times 10^{-4} h^{-3} \text{Mpc}^{-3}$
"SAG-dens n_{CMASS} "

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SAG

Why?!

Galaxy Assembly Bias in SAMs



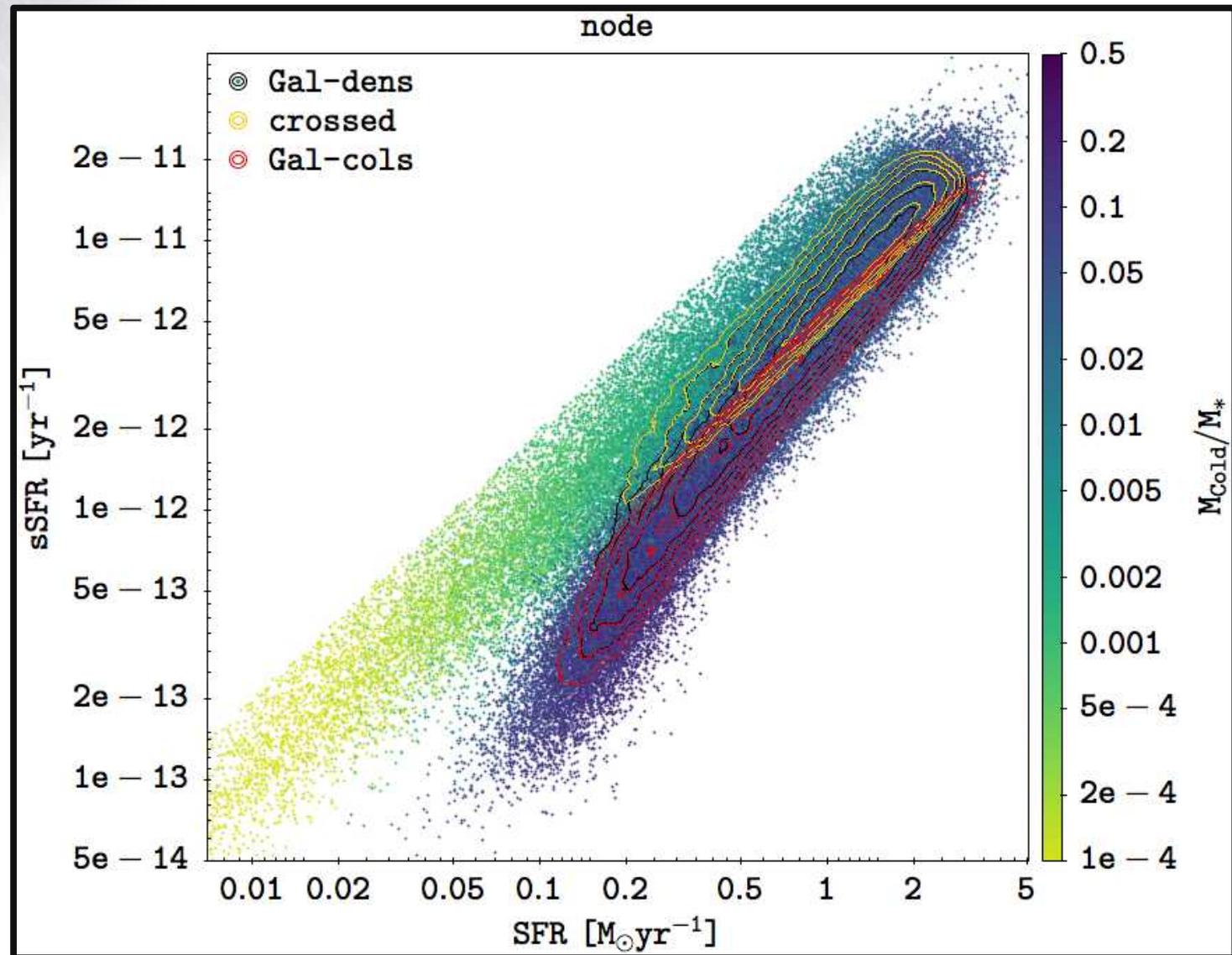
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Observations	<p>Cooper et al. (2010, SDSS NYU-VAGC, significant bias on red sequence) Lacerna et al. (2014, small on centrals) Gallart et al. (2015, local group, dwarf galaxies), Lin et al. (2015, no signal) Guo et al. (2015, BOSS-CMASS), Zentner et al. (2016) on cluster Miyatake et al. (2016); Zu et al. (2017) Dvornik et al. (2017, no signal of halo assembly bias in GAMA) Montero-Dorta et al. (2017, BOSS-CMASS SF-history), Tinker et al. (2018) small: Zu & Mandelbaum (2016); Niemiec et al. (2018); Xu & Zheng (2018)</p>
Simulations	
Dark Matter	<p>Artale, Pedrosa, et al. (2019) Salcedo et al. (2018, Las Damas) Vakili & Hahn (2016, sMDPL) Sunayama et al. (2016, Vmax scale dependency, Bolshoi, and MD) Chue et al. (2018, in massive halos MD), Sato-Polito et al. (2018, comparison MDs)</p>
Hydro	<p>Bray et al. (2016, Illustris, galactic conformity), Xu & Zheng (2018, Illustris) Chaves-Montero et al. (2016); Artale et al. (2018, Eagle)</p>
HOD and HAM	<p>Zentner et al. (2014, assembly bias in HOD are highly affected by systematic errors) Saito et al. (2016, BOSS-CMASS) Paranjape and Padmanabhan (2017, Separate Universe-technique) Zehavi et al. (2018, occupation variation)</p>
SAMs	<p>Lacerna & Padilla (2011, SAGv2 by Lagos, Cora, and Padilla (2008)), Wang et al. (2013) Jung et al. (2014, ySAM, small galaxy assembly bias) Pujol & Gaztañaga (2014, small galaxy bias, at low Mhalo halo bias take over) Tojeiro et al. (2017, LGALAXIES and GAMA, halo assembly bias) Lacerna et al. (2018, LGALAXIES and Galform) Padilla et al. (2018, LGALAXIES, BOSS-CMASS, RSD) Raouf et al. (2018, Radio-SAGE, major merger, clustering, halo mass assembly) Contreras et al. (2019, LGALAXIES, redshift evolution)</p>

Why?!

Environmental Dependency

Vweb Cui+19

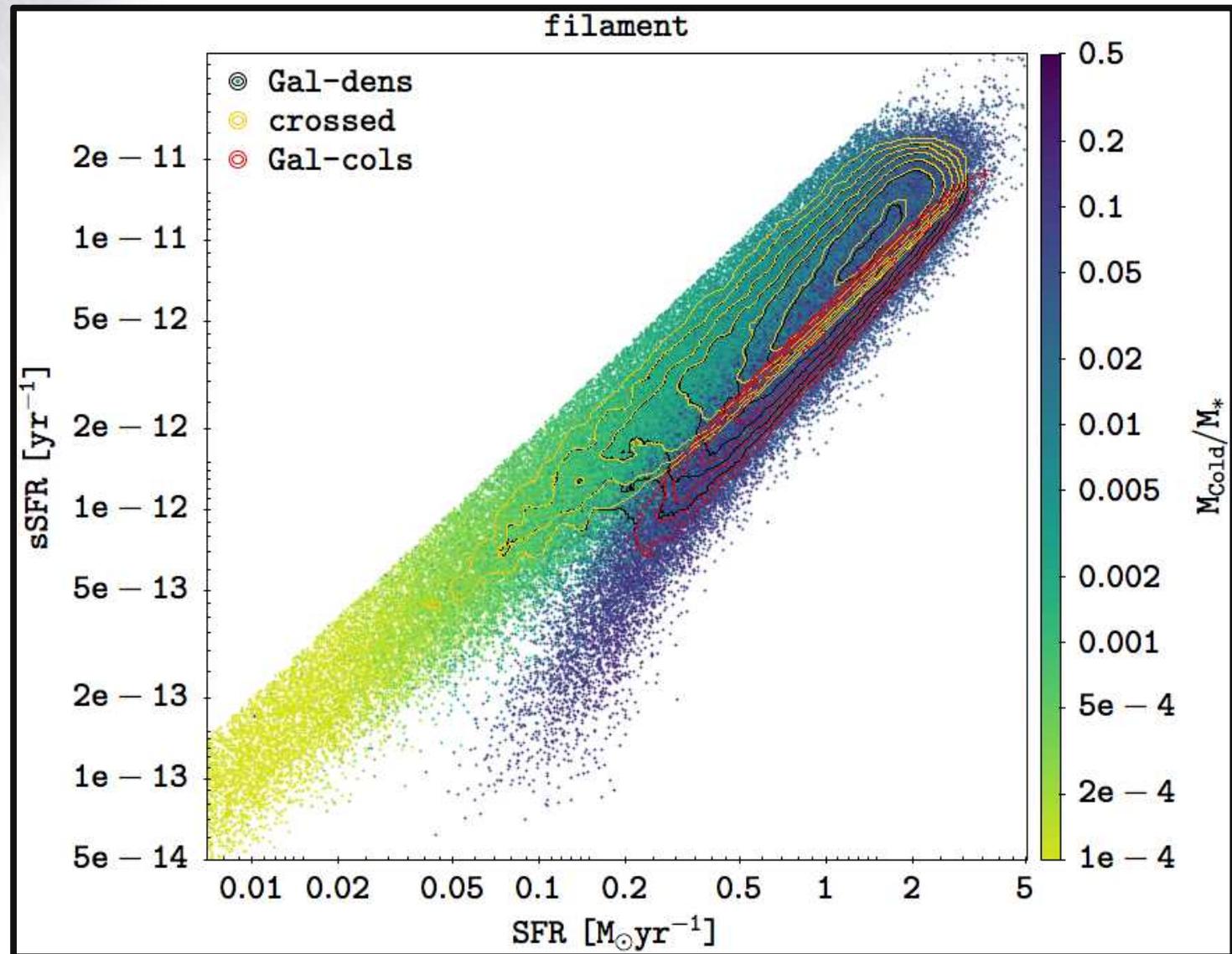


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Why?!

Environmental Dependency

Vweb Cui+19



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Why?!

Environmental Dependency

sample name & population	environment	fraction of galaxies	$\log_{10}(M_{200c} [M_{\odot}])$	$\log_{10}(M_{*} [M_{\odot}])$	$\log_{10}(sSFR [\text{yr}^{-1}])$	Z_{Cold}
Gal-cols	knot	0.61	$13.79^{+0.21}_{-0.19}$	$11.44^{+0.11}_{-0.09}$	$-11.77^{+0.35}_{-0.37}$	$9.10^{+0.18}_{-0.16}$
Gal-cols	filament	0.37	$13.57^{+0.19}_{-0.17}$	$11.37^{+0.09}_{-0.05}$	$-11.52^{+0.30}_{-0.34}$	$9.18^{+0.21}_{-0.17}$
Gal-dens	knot	0.52	$13.55^{+0.27}_{-0.28}$	$11.28^{+0.15}_{-0.14}$	$-11.53^{+0.36}_{-0.42}$	$9.25^{+0.36}_{-0.23}$
Gal-dens	filament	0.41	$13.22^{+0.23}_{-0.26}$	$11.14^{+0.13}_{-0.14}$	$-11.36^{+0.32}_{-0.44}$	$9.55^{+0.30}_{-0.35}$
Gal-dens Pop (A)	knot	0.26	$13.13^{+0.32}_{-0.34}$	$11.05^{+0.07}_{-0.09}$	$-11.40^{+0.38}_{-0.53}$	$9.76^{+0.19}_{-0.29}$
Gal-dens Pop (A)	filament	0.62	$13.01^{+0.22}_{-0.22}$	$11.02^{+0.08}_{-0.11}$	$-11.40^{+0.39}_{-0.58}$	$9.80^{+0.19}_{-0.25}$
Gal-dens Pop (B)	knot	0.54	$13.69^{+0.23}_{-0.21}$	$11.36^{+0.13}_{-0.09}$	$-11.58^{+0.35}_{-0.40}$	$9.13^{+0.21}_{-0.17}$
Gal-dens Pop (B)	filament	0.44	$13.46^{+0.18}_{-0.16}$	$11.29^{+0.09}_{-0.07}$	$-11.33^{+0.26}_{-0.32}$	$9.23^{+0.23}_{-0.19}$
(i)	(ii)	(iii)	(iv)	(v)	(vi)	(vii)

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Why?!

Environmental Dependency

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Why?!

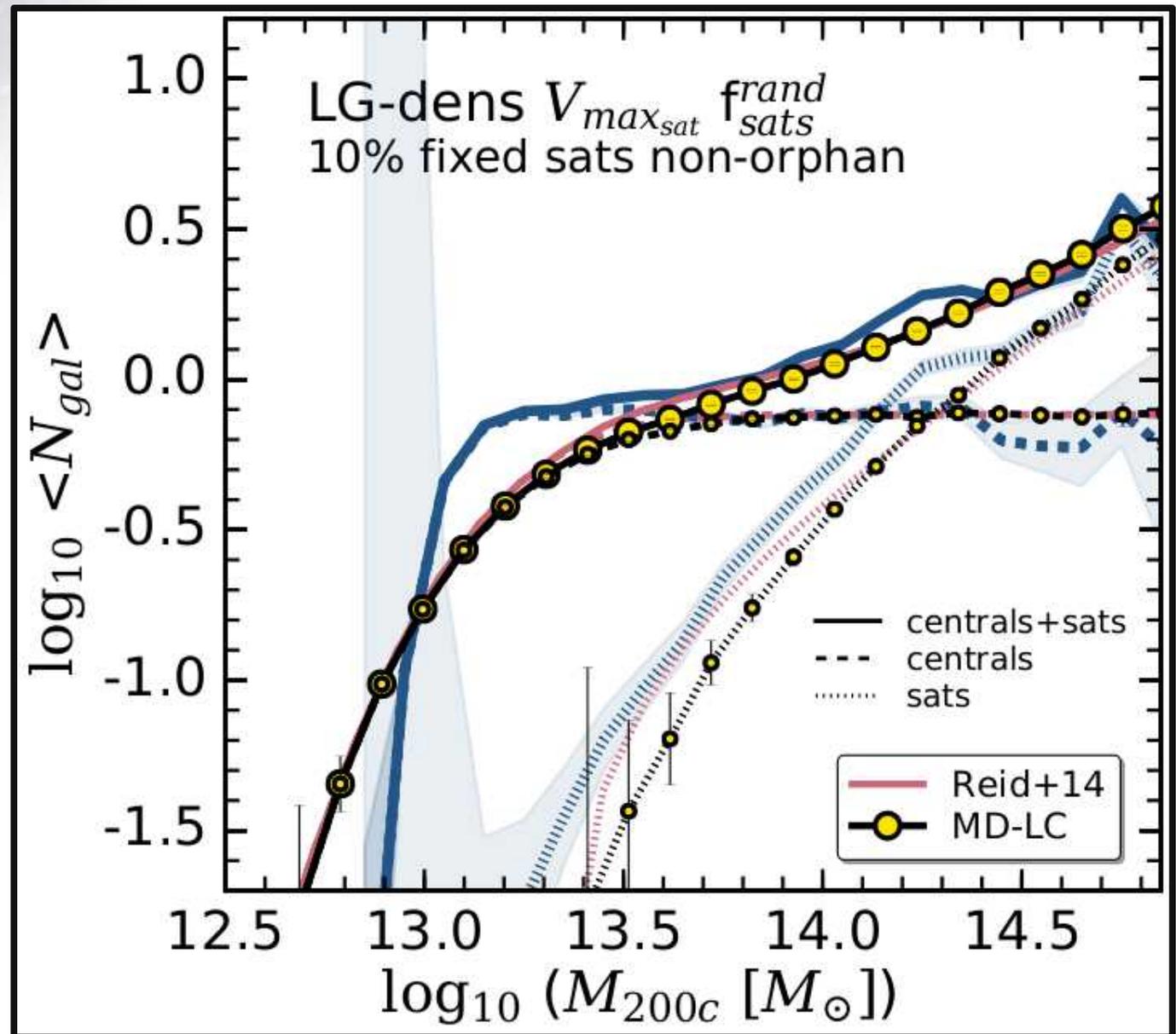
Environmental Dependency

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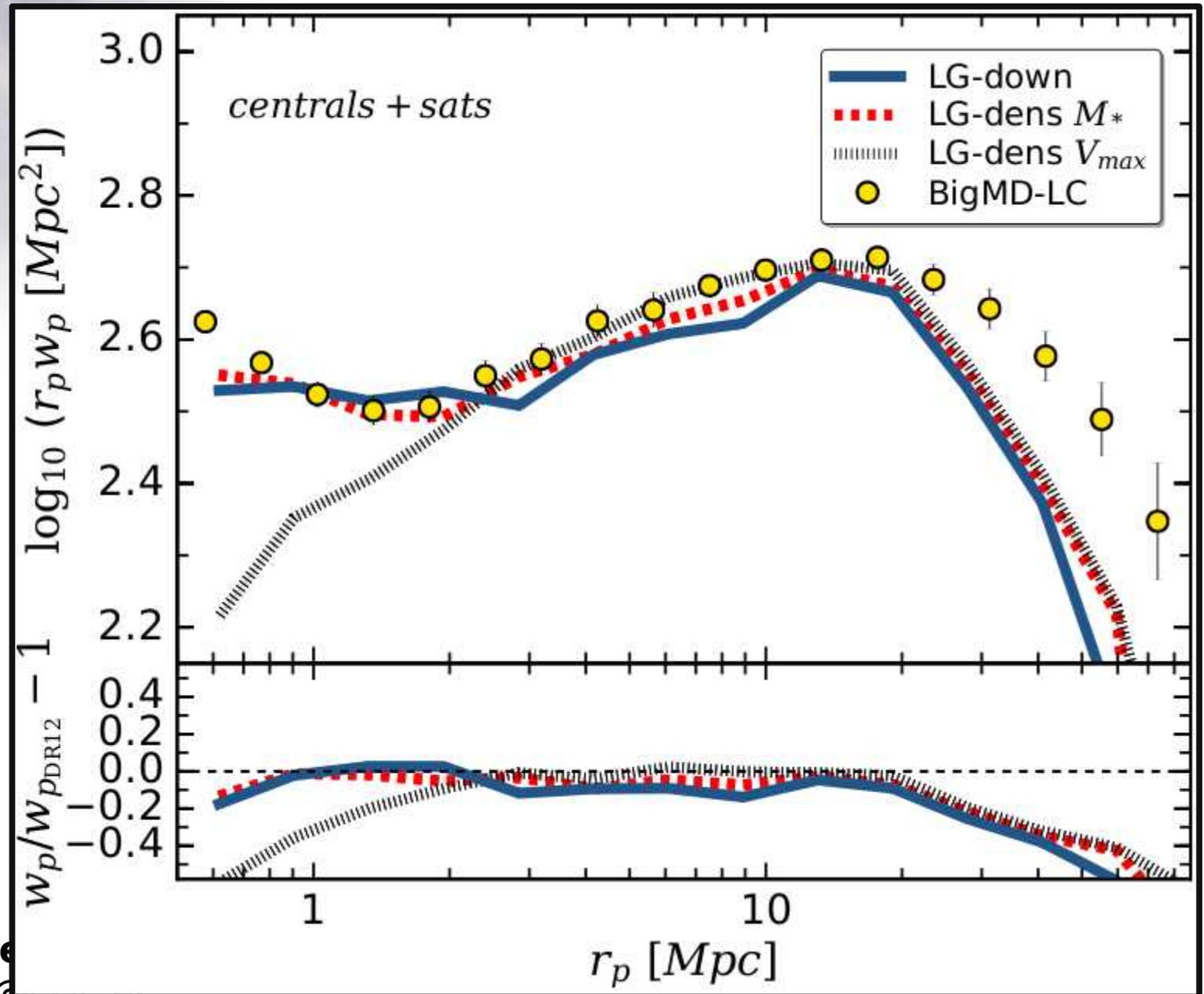
L-GALAXIES Guo+13

Outlook



L-GALAXIES Guo+13

Outlook



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Conclusions

Who?!

CMASS hard to model with SAMs – alternatives: density samples / down sampling

What?!

Dependencies of certain properties on environment: halo mass / star forming properties

Why?!

Could provide tracers for the galaxy assembly bias! Stay tuned for further results ...

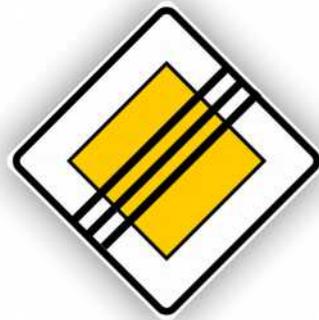
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Collaborations and Credits

Knebe+18

arXiv: 1710.08150



Stoppacher+19

arXiv: 1902.05496



F. Prada, A. Montero-Dorta, S. Rodriguez-Torres,
G. Favole, A. Knebe, W. Cui,
A. Orsi, S. Contreras, A. Benson, C. Behrens,
S. Cora, C. Vega, V. Gonzalez-Perez,
D. Croton, A. Stevens, M.L.L. Dantas ++

Image credits title page: <https://www.imcdb.org/v061628.html>

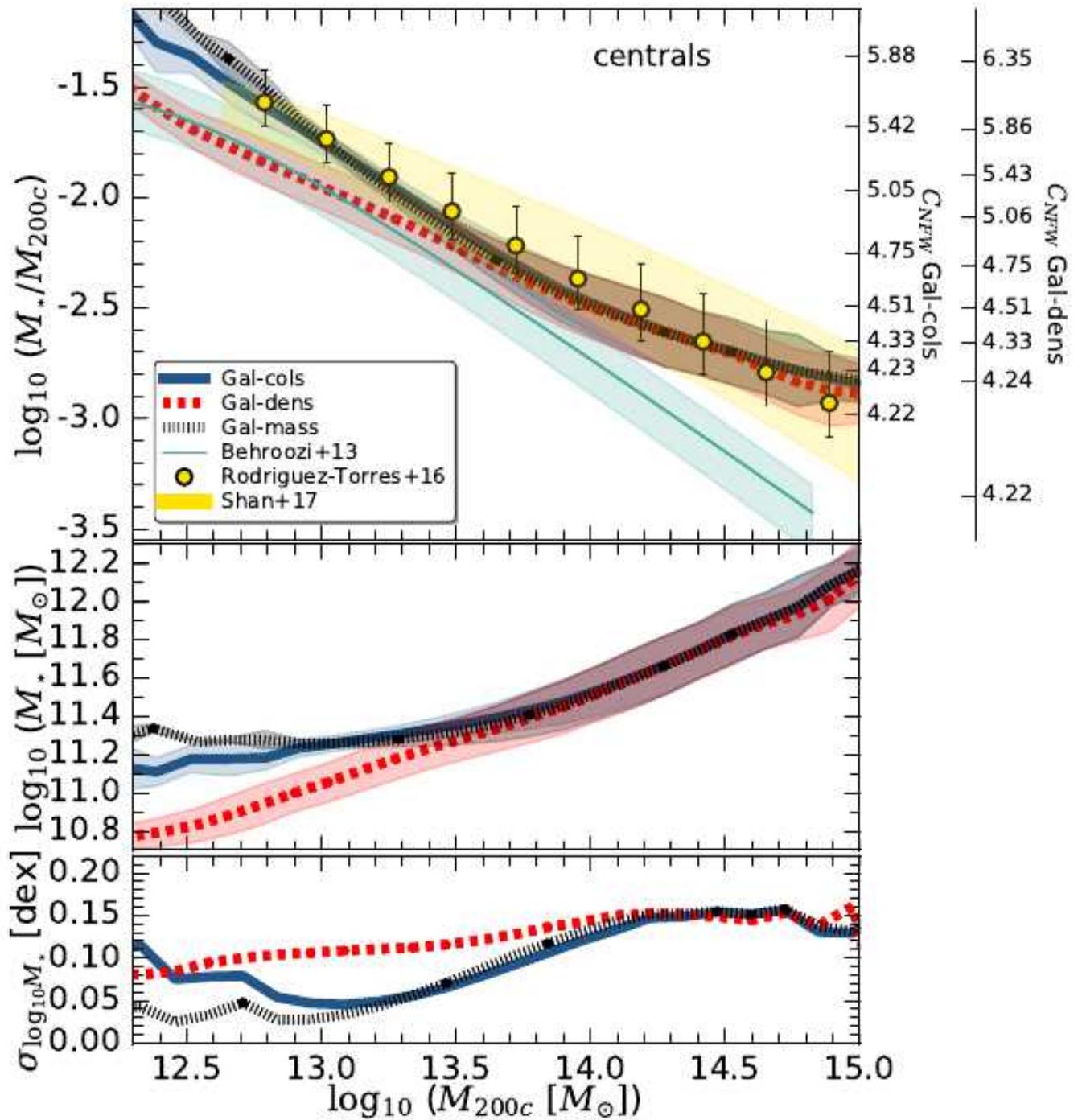
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Extra-Slides

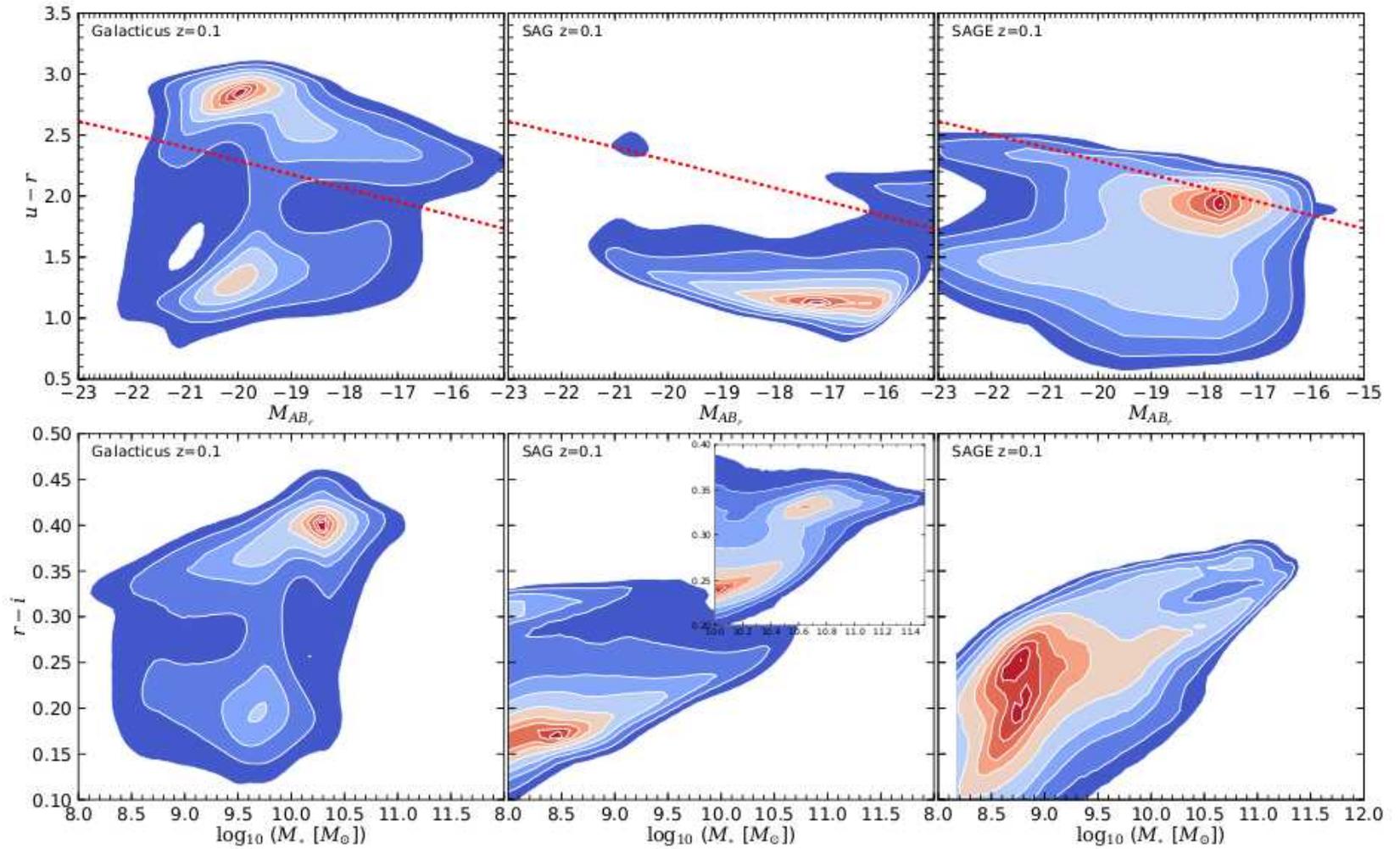
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Luminosities



Knebe et al. 2018 (1710.08150v1)